

# Physics on the boundary between Quantum Mechanics and Gravity

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**Quantum Fields, Gravity and Information**

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Why introduce something new into Quantum mechanics?

- Schrödinger Cat States
- Description of measurement process
- Linearity of quantum mechanics vs Nonlinearity of General Relativity

## Measurement in quantum mechanics

### Density Matrix

$$\rho \text{ before measurement} \rightarrow \frac{P_n \rho P_n}{\text{Tr}(P_n \rho P_n)} \text{ after measurement}$$

### Measurement Problem

$$|s_n\rangle \otimes |A_1\rangle \rightarrow |s_n\rangle \otimes |A_n\rangle$$

Problem:

$$\frac{1}{\sqrt{2}} (|s_n\rangle + |s_l\rangle) \rightarrow \frac{1}{\sqrt{2}} (|s_n\rangle \otimes |A_n\rangle + |s_l\rangle \otimes |A_l\rangle)$$

Entanglement between Apparatus and measured system.

Interpretation:

Measuring either  $|s_n\rangle \otimes |A_n\rangle$  or  $|s_l\rangle \otimes |A_l\rangle$  with probability  $\frac{1}{2}$ .

This leads to two different time evolutions:

- Schrödinger evolution  $\rightarrow$  linear, deterministic
- Wavepacket reduction  $\rightarrow$  nonlinear, stochastic

## Possible ways to deal with this problem[Bassi,2003]:

- Incompleteness
- formal Completeness
- with different Individuals
- with identical Individuals
- Two dynamical principles
- Unifyied dynamics

Dynamical reduction models:

Simple example:

$$|\psi\rangle = \frac{1}{\sqrt{2}}(|x_1\rangle + |x_2\rangle)$$

$$|\psi\rangle\langle\psi| = \frac{1}{2} \begin{pmatrix} 1 & 1 \\ 1 & 1 \end{pmatrix}$$

Look for "collapsed" system

$$\frac{1}{2} \begin{pmatrix} 1 & 1 \\ 1 & 1 \end{pmatrix} \rightarrow \frac{1}{2} \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}$$

Also evolution has to give the right mixture, i.e. individual parts have to have definite macroscopic properties.

## Linear, stochastic

$$i\hbar\partial_t|\psi(t)\rangle = \sum_i P_i V_i(t)|\psi(t)\rangle$$

$$\langle\langle V_i(t)\rangle\rangle = 0$$

$$\langle\langle V_i(t)V_j(t')\rangle\rangle = \gamma\delta_{ij}\delta(t-t')$$

$$\partial_t\langle x_i|\rho(t)|x_j\rangle = \gamma(\delta_{ij} - 1)\langle x_i|\rho(t)|x_j\rangle$$

BUT

$$\langle\psi(t)|P_i|\psi(t)\rangle = \langle\psi(0)|P_i|\psi(0)\rangle$$

## Nonlinear, deterministic

Ensembles  $E_1, E_2$

$$E_1 \{a_i, |\phi_i\rangle\} \quad E_2 \{b_i, |\varphi_i\rangle\}$$

Purification

$$\sum_i a_i |\phi_i\rangle \otimes |\eta_i\rangle = \sum_i b_i |\varphi_i\rangle \otimes |\zeta_i\rangle$$

Two regions  $X_1 \rightarrow \phi, \varphi, X_2 \rightarrow \eta, \zeta$   
 Create ensembles by measuring in  $X_2$ . Due to No-Signalling  $\rightarrow$  only linear operations allowed! [Gisin, 1989]



Collapse Models are Nonlinear and Stochastic.

Bassi, Dürr, Hinrichs 2013:[Bassi,2013]

"Collapse models are the only possible nonlinear extensions of the Schrödinger equation, compatible with the no-faster-than-light assumption."

Only equations of type

$$d\psi_t = \left( -iHdt + \sum_{k=1}^n (L_k - l_{k,t})dW_{k,t} - \frac{1}{2} \sum_{k=1}^n (L_k^\dagger L_k + 2l_{k,t}L_k + |l_{k,t}|^2)dt \right) \psi_t$$

$$l_{k,t} = \frac{1}{2} \langle \psi_t | (L_k^\dagger + L_k) | \psi_t \rangle$$

are allowed.

## The CSL Model

$$\partial_t \rho(t) = -\frac{i}{\hbar} [H, \rho(t)] + \gamma \int d^3x N(x) \rho(t) N(x) - \frac{\gamma}{2} \int d^3x \{N^2(x), \rho(t)\}$$

where  $N(x)$  denotes the preferred basis

$$N(x) = \int d^3y g(y-x) a^\dagger(y) a(y) \quad g(x) = \left(\frac{\alpha}{2\pi}\right)^{\frac{3}{2}} e^{-\frac{\alpha}{2}(x)^2}$$

## Simple model:

$$\partial_t \rho(t) = \gamma \left(\frac{\alpha}{4\pi}\right)^{\frac{3}{2}} \sum_k \left[ \sum_i N_i^{(k)} \rho(t) N_i^{(k)} - \frac{1}{2} \sum_i \{N_i^{(k)2}, \rho(t)\} \right]$$

$$\langle n_1^{(k)}, \dots | \rho(t) | m_1^{(k)}, \dots \rangle = e^{-\frac{\gamma}{2} \left(\frac{\alpha}{4\pi}\right)^{\frac{3}{2}} \sum_{i,k} (n_i^{(k)} - m_i^{(k)})} \langle n_1^{(k)}, \dots | \rho(0) | m_1^{(k)}, \dots \rangle$$

Question:

Meaning of Parameters?

Replace Number operator with Mass Density operator

$$M(x) = \sum_k m_k N_k(x)$$

$$\partial_t \rho(t) = -\frac{i}{\hbar} [H, \rho(t)] + \frac{\gamma}{m_0^2} \int d^3x M(x) \rho(t) M(x) - \frac{\gamma}{2m_0^2} \int d^3x \{M^2(x), \rho(t)\}$$

Thus relate collapse to the mass density of the particles. Still two arbitrary parameters needed.

One step further:

Diosi's proposition of collapse due to gravity.

(QMUDL: quantum mechanics with universal density localization)

Introduce mass density

$$f(x) = \frac{M}{V} \theta(R - |x' - x|)$$

$$\partial_t \rho(t) = -\frac{i}{\hbar} [H, \rho(t)] - \frac{G}{2\hbar} \int \int \frac{d^3 x_1 d^3 x_2}{|x_1 - x_2|} [f(x_1), [f(x_2), \rho(t)]]$$

for description of decoherence processes due to gravity.

## Assume Gravitational Potential as cause for Collapse

$$U(x) = -G \int \frac{f(x_1)f(x_2)}{|x_1 - x_2|} dx_1 dx_2$$

This leads to

$$\langle x' | \rho(t) | x'' \rangle = e^{\Gamma(|x' - x''|)t} \langle x' | \rho(0) | x'' \rangle$$

where

$$\Gamma(|x' - x''|) = \frac{1}{\hbar} (U(0) - U(|x' - x''|))$$

Assume Gravitational Potential as cause for Localization

Modify Schrödinger Equation:

$$i\hbar\partial_t\psi = \left( -\frac{\hbar^2}{2m}\Delta - Gm^2 \int \frac{|\psi^2|}{|x-y|} d^3y \right) \psi$$

$$\Delta\Phi = 4\pi Gm|\psi|^2$$

**Nonlinear Integro-Differential Equation!**

## Scaling of Schrödinger Newton

$$m_\mu = \mu m$$

$$t_\mu = \mu^5 t$$

$$r_\mu = \mu^3 r$$

$$\psi_\mu = \mu^{\frac{9}{2}} \psi$$

# Numerical solutions: Solitons!

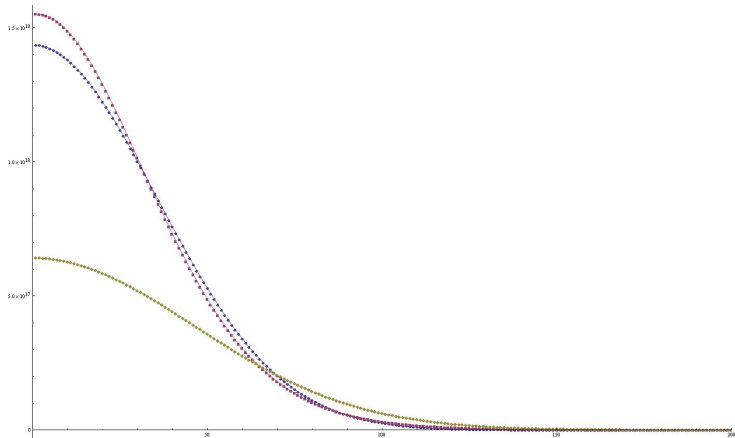


Abbildung : Solitonic solution of the Schrödinger Newton equation. Stepsize:  $r = 10^{-8}m$ , Blue: Gauss Packet at  $t = 0$ , Red: Schrödinger Newton solution after 20000s, Yellow: Schrödinger solution after 20000s.



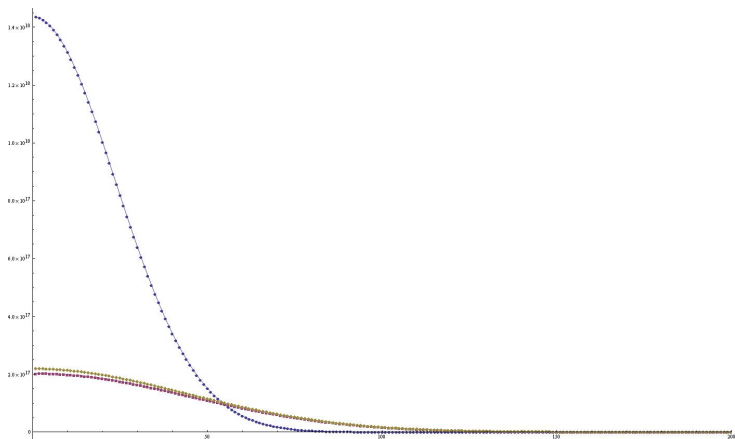


Abbildung : Solution of the Schrödinger Newton equation. Stepsize:  $r = 10^{-14} m$ ,  
Blue: Gauss Packet at  $t = 0$ , Red: Schrödinger Newton solution after  $10^{-10} s$ ,  
Yellow: Schrödinger solution after  $10^{-10} s$ .

## Conclusions

- Framework beyond quantum mechanics.
- Is the only feasible Ansatz that does not contradict the No-faster-than-light assumption.
- Allows explanation of the ad-hoc collapse of the wavefunction without introducing a second time evolution.
- Introduced parameters may be linked to Gravity.
- Experimental tests may be possible to distinguish these models from standard quantum mechanics.

